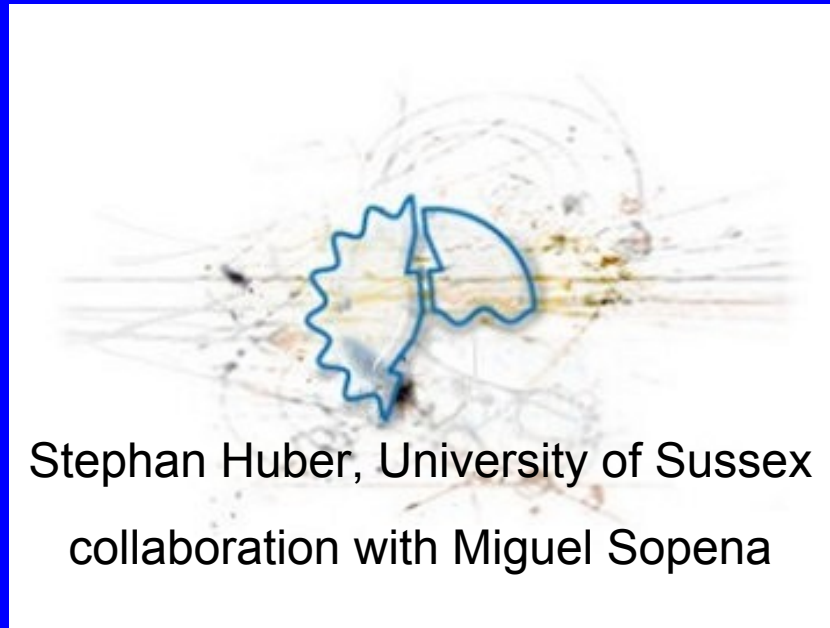


How fast do (electroweak) bubbles move?



Pascos 09, Hamburg, Desy, July 2009

Relics of the phase transition

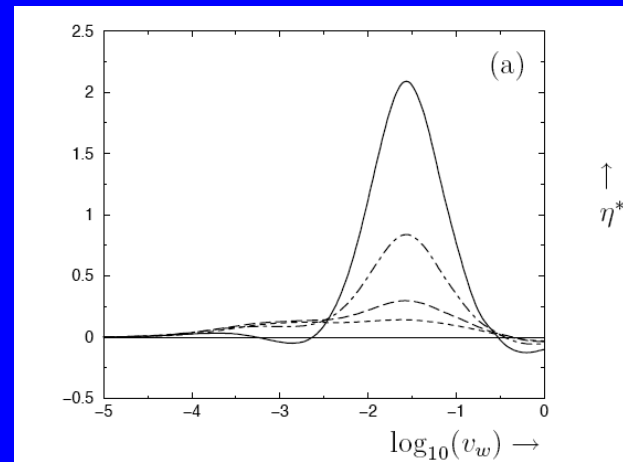
- gravitational waves

dependence on the wall velocity!

$$h_0^2 \Omega_{\text{det}} \simeq 1.2 \times 10^{-7} \kappa^2 \langle R \rangle^2 H_*^2 \left[\frac{\alpha}{\alpha + 1} \right]^2 \left[\frac{v_b^2}{0.24 + v_b^3} \right]^2 \left[\frac{100}{g_*} \right]^{1/3}.$$

Kamionkowski, Kosowsky, Turner '93

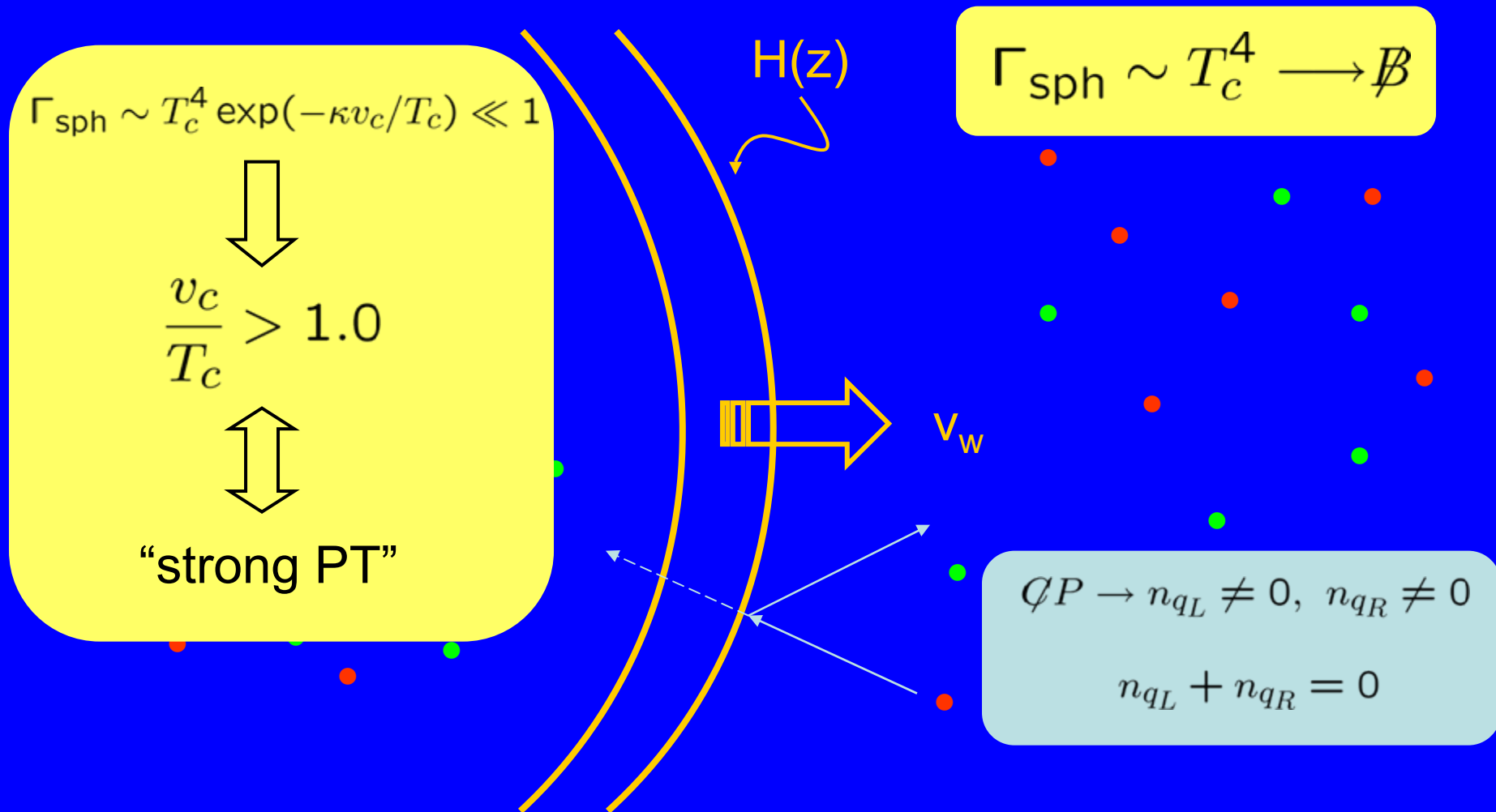
- baryon asymmetry



SH, Schmidt '00

- magnetic fields

Electroweak baryogenesis



1) The strength of the PT

Thermal potential (or lattice):

$$V(H, T) = m^2(T)H^2 - E(T)H^3 + \lambda(T)H^4$$

- Boson loops:

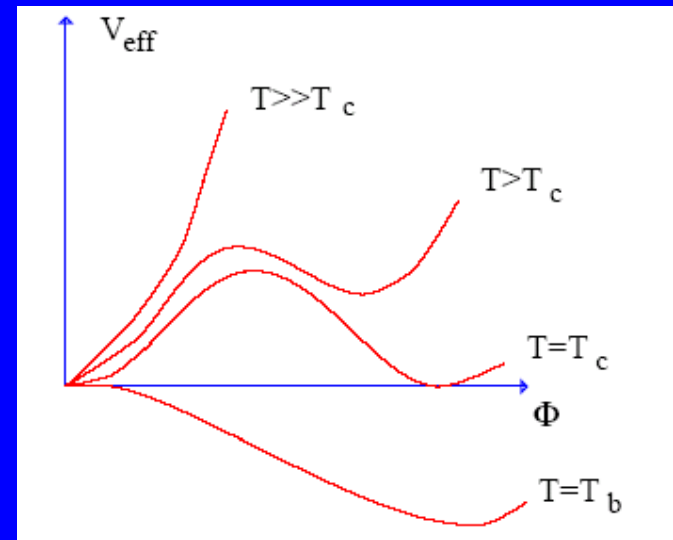
SM: gauge bosons

SUSY: light stops

2HDM: heavy Higgses

- tree-level: **extra singlets**: λSH^2 , NMSSM, etc.

- **replace H^4 by H^6** , etc.



SM + higher-dim. operators

$$V(H) = -\mu^2 |H|^2 + \lambda |H|^4 + \frac{1}{M^2} |H|^6$$

Zhang '93

Grojean et al. '04

maybe related to strong dynamics at the TeV scale, such as technicolor or gravity?
(or simply comes from integrating out extra scalars)

two parameters, $(\lambda, M) \leftrightarrow (m_h, M)$

λ can be negative \rightarrow bump because of $|H|^4$ and $|H|^6$

$$\begin{aligned} V_{\text{eff}}(\phi, T) = & \frac{1}{2} \left(-\mu^2 + \left(\frac{1}{2}\lambda + \frac{3}{16}g_2^2 + \frac{1}{16}g_1^2 + \frac{1}{4}y_t^2 \right) T^2 \right) \phi^2 \\ & - \frac{g_2^3}{16\pi} T \phi^3 + \frac{\lambda}{4} \phi^4 + \frac{3}{64\pi^2} y_t^4 \phi^4 \ln \left(\frac{Q^2}{c_F T^2} \right) \\ & + \frac{1}{8M^2} (\phi^6 + 2\phi^4 T^2 + \phi^2 T^4). \end{aligned}$$

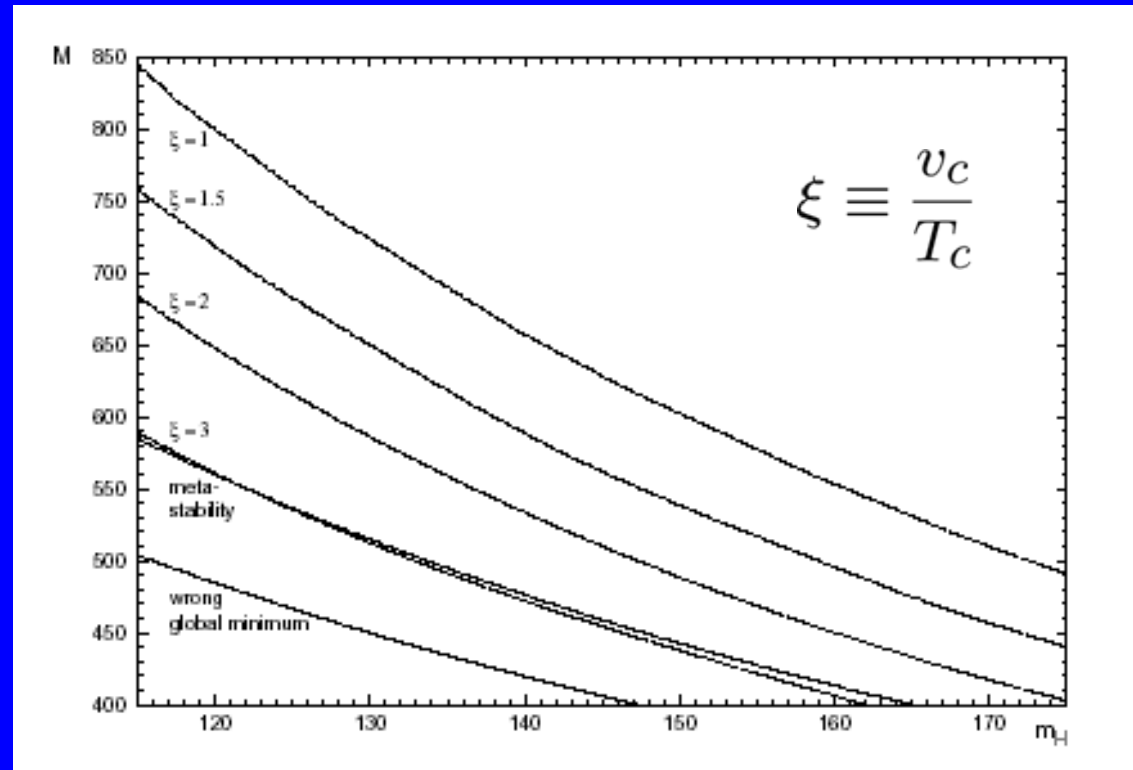
Results for the PT

Evaluating the 1-loop
thermal potential:

strong phase transition
for $M < 850$ GeV
up to $m_h \sim 170$ GeV

(LEP bound applies,
 $m_h > 114$ GeV)

wall thickness
 $2 < L_w T_c < 16$



Bödeker, Fromme, S.H., Seniuch '04

Similar results, including Higgs cubic terms

Delaunay, Grojean, Wells '07

2) Dynamics of the transition

At the critical temperature T_c the two minima are degenerate

Bubble nucleation starts at $T < T_c$ with a **rate**

$$\Gamma = A T^4 e^{-S_3/T},$$

Where the **bubble energy** is

$$S_3 = 4\pi \int dr r^2 \left[\frac{1}{2} \left(\frac{d\phi}{dr} \right)^2 + V(\phi, T) \right]$$

The **bubble configuration** follows from

(with appropriate BC's)

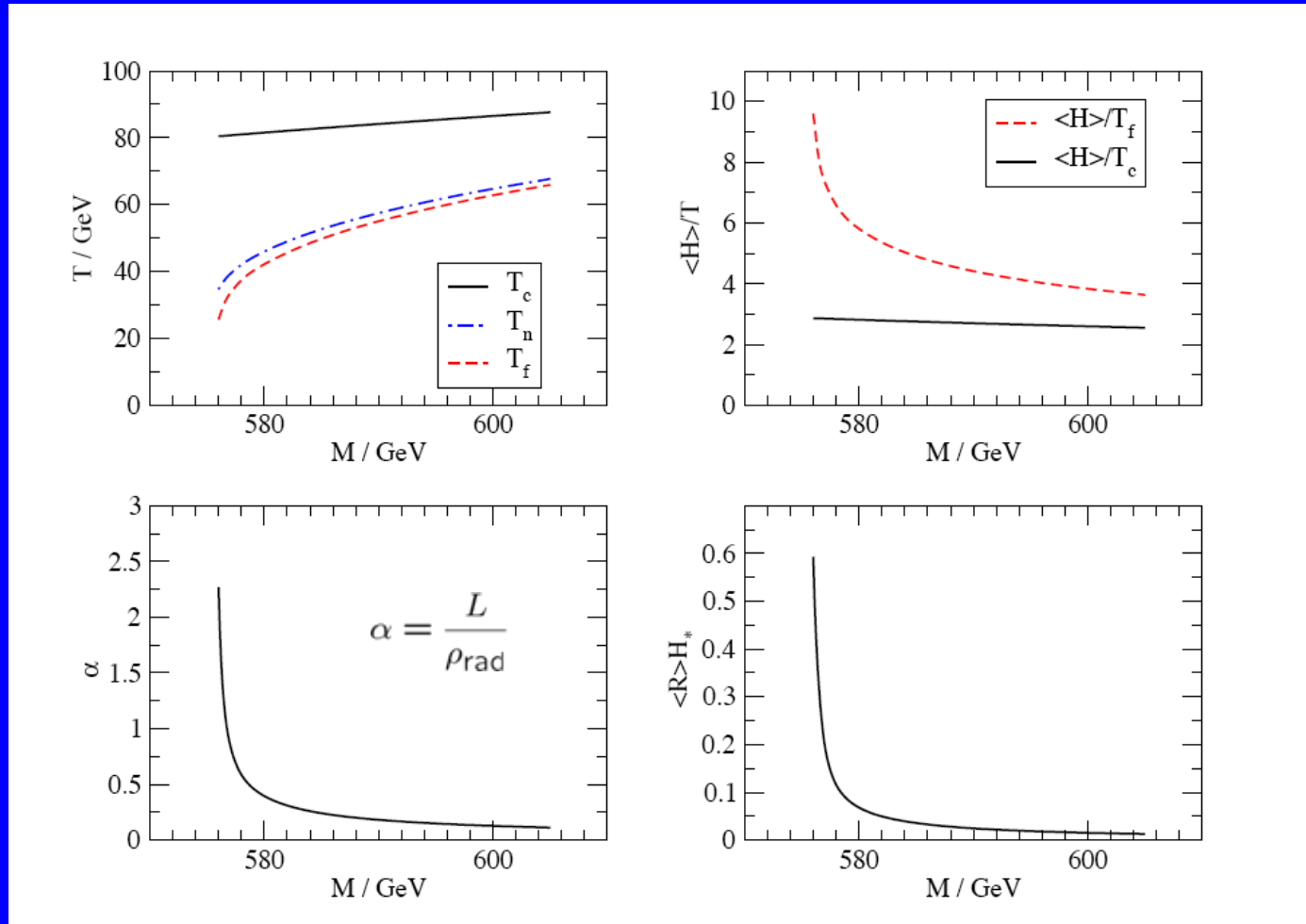
$$\frac{\partial^2 \phi}{\partial r^2} + \frac{2}{r} \frac{\partial \phi}{\partial r} = V'(\phi)$$

This bounce solution is a **saddle point**, not a minimum

→ difficult to compute for multi field models (one field: shooting)

For a **algorithm** see Konstandin, S.H. '06

Key parameters of the phase transition: Φ^6 model, $m_h=120$ GeV



S. H. &
Konstandin '07

Compute as function of temperature: bubble configurations \rightarrow E

$$V(H) = -\mu^2 |H|^2 + \lambda |H|^4 + \frac{1}{M^2} |H|^6$$

nucleation rate $\Gamma \sim \exp(-E)$
bubble distribution $\rightarrow R$

The wall velocity:

Friction with the plasma balances the pressure

Distinguish: supersonic vs. subsonic ($v_s^2=1/3$)

Standard model: $v_b \sim -0.35 - 0.45$ for low Higgs masses [Moore, Prokopec '95]

MSSM: $v_b \sim 0.05$ [John, Schmidt '00]

All other models: no detailed computations

For very strong phase transitions: **bubbles become supersonic**, velocity dominated by hydrodynamics (neglect friction) [Steinhardt '82]

$$v_b(\alpha) = \frac{1/\sqrt{3} + \sqrt{\alpha^2 + 2\alpha/3}}{1 + \alpha}$$

(for sufficiently large $\alpha > \text{few } \%$?)

When does this fail ??

**Recently: can the walls run away? [Bodeker, Moore '09]

How to compute the wall velocity?

Main ingredients: pressure difference vs. plasma friction

Also important: reheating due to release of latent heat

Microscopic description: Moore, Prokopec '95

$$(\partial_t + \dot{z}\partial_z + \dot{p}_z\partial_{p_z})f = C[f]$$

$$\square\phi + V_T'(\phi) + \sum \frac{dm^2}{d\phi} \int \frac{d^3p}{(2\pi)^3} \frac{1}{2E} \delta f(p, x) = 0$$

$$f = \frac{1}{1 + \exp\frac{E - E\delta T/T - p_z v - \mu}{T}}$$

(fluid ansatz)

$$\begin{aligned} av_w \frac{\mu'}{T} + v_w \frac{\delta T'}{T} + \frac{1}{3}v' + F_1 &= -\Gamma_{\mu 1} \frac{\mu}{T} - \Gamma_{T1} \frac{\delta T}{T} \\ bv_w \frac{\mu'}{T} + v_w \frac{\delta T'}{T} + \frac{1}{3}v' + F_2 &= -\Gamma_{\mu 2} \frac{\mu}{T} - \Gamma_{T2} \frac{\delta T}{T} \\ b \frac{\mu'}{T} + \frac{\delta T'}{T} + v_w v' + 0 &= -\Gamma_v v \end{aligned}$$

$$F_1 = -\frac{v_w \ln 2}{9\zeta_3} \frac{(m^2)'}{T^2}, \quad F_2 = -\frac{v_w \zeta_2}{42\zeta_4} \frac{(m^2)'}{T^2}$$

(force terms)

$$(F = \dot{p}_z = -\partial_z E(z, p_z))$$

→ Complicated set of coupled field equations

and Boltzmann equations

need many scattering rates

SM: $v \sim 0.35 - 0.45$

Simplified approach: (Ignatius, Kajantie, Kurki-Suonio, Laine '94)

1) describe friction by a friction coefficient $1/\Gamma$

2) Model the fluid by a fluid velocity and temperature

$$\phi''(x) = \frac{\partial V}{\partial \phi} + \frac{v\gamma}{\Gamma} \phi'(x)$$

← 1 + 1 dimensions

$$(4aT^4 - T \frac{\partial V}{\partial T}) \gamma^2 v = \text{const.}$$

$$(4aT^4 - T \frac{\partial V}{\partial T}) \gamma^2 v^2 + aT^4 + \frac{1}{2} \phi'(x)^2 - V = \text{const.}$$

3) Determine Γ from fitting the to the full result by Moore and Prokopec

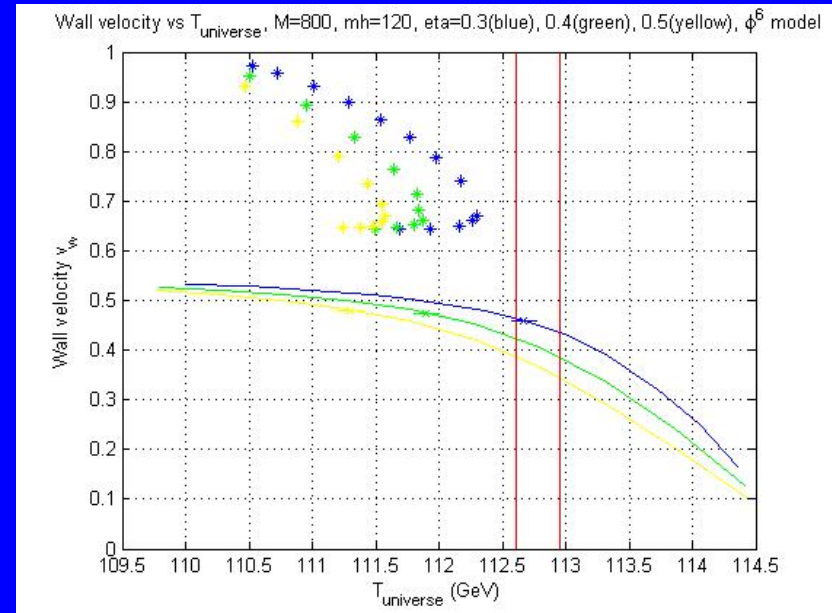
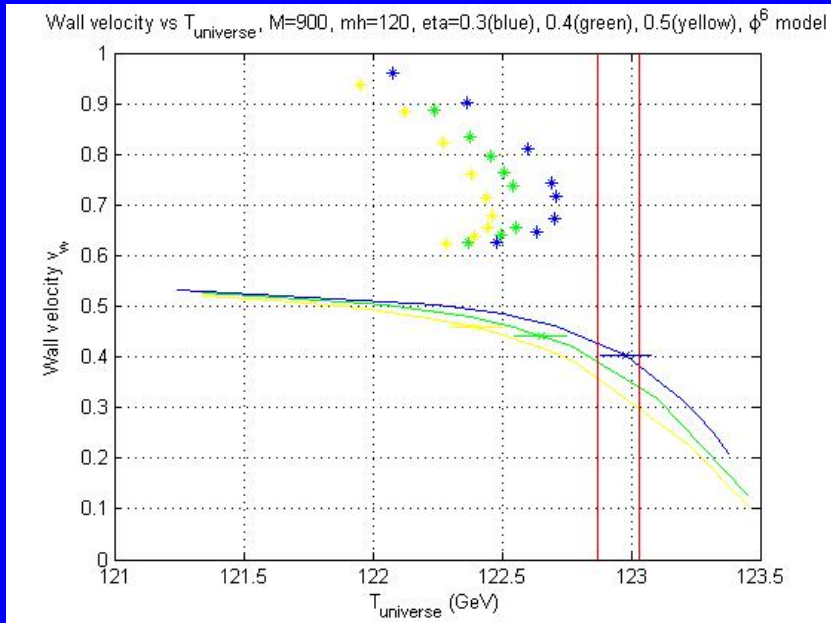
(with Miguel Sopena, see also Megevand, Sanchez '09)

We find: a **good fit** with a universal Γ is possible

→ the formalism should describe situations with SM friction well

→ study models with SM friction, but different potential, e.g. ϕ^6 model

Results for Φ^6 model



- for a moderately strong phase transition, the wall moves subsonic in the dim-6 model (good for baryon asymmetry, bad for GW's)
- instable bubbles for some temperatures?

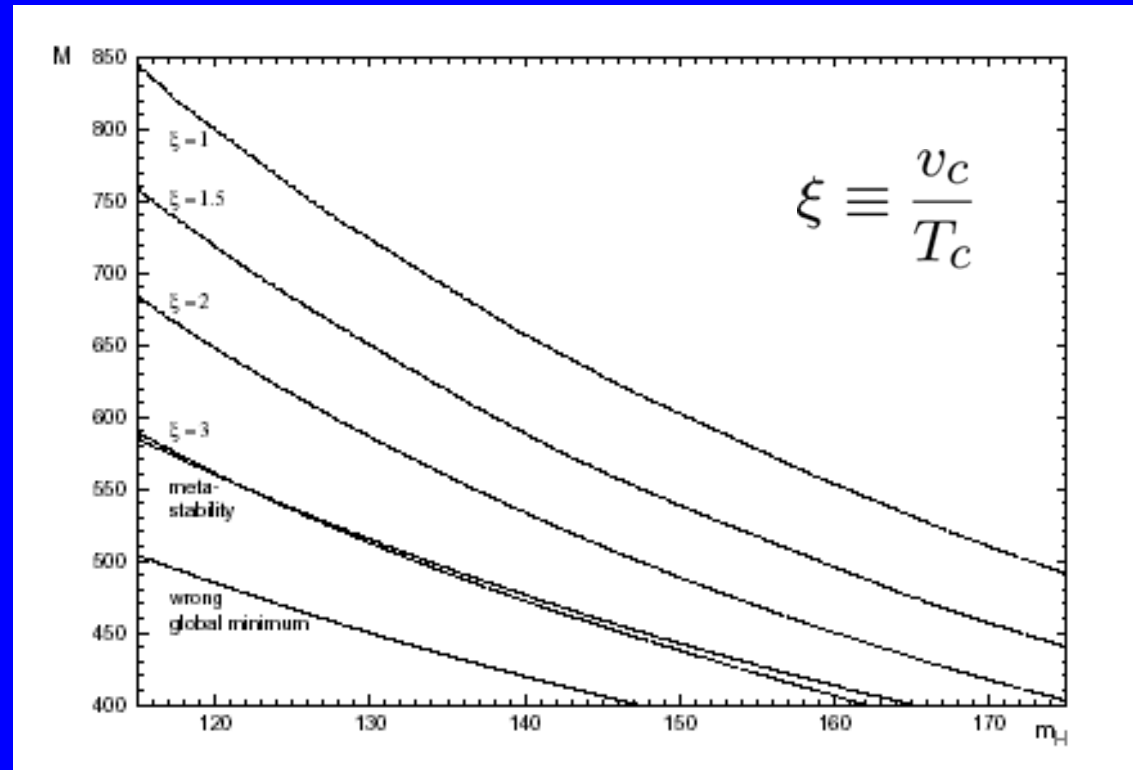
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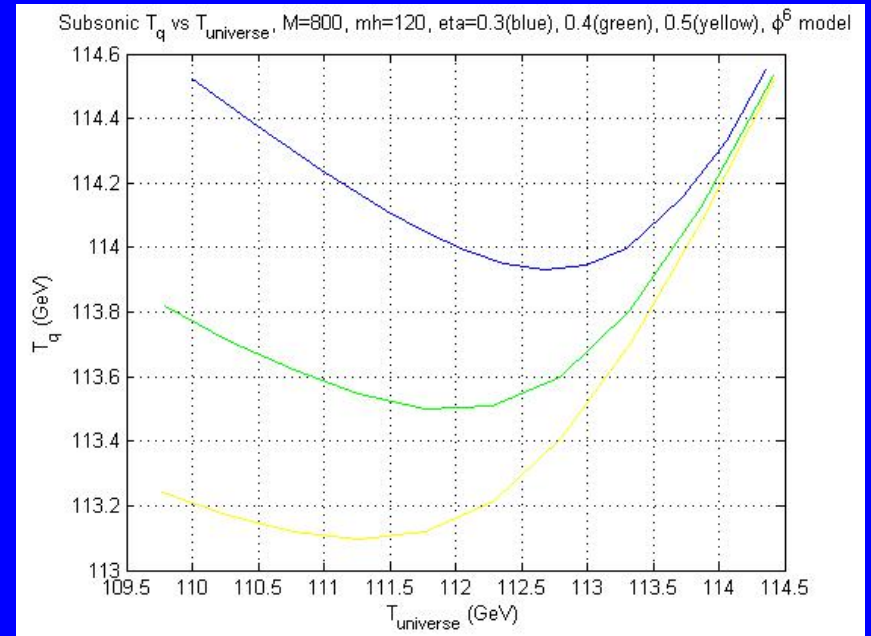
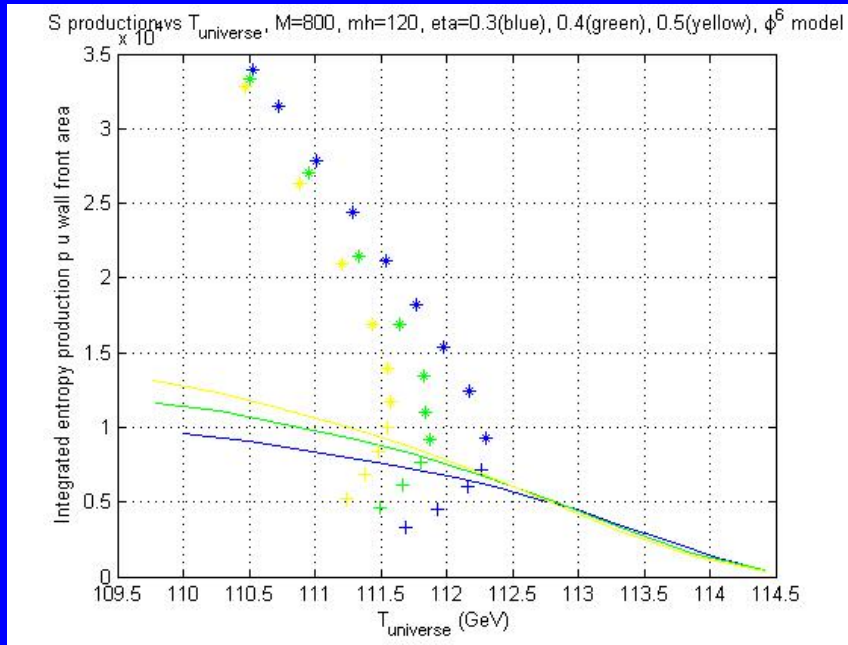
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Detonations, if possible, give the biggest entropy

Instability due to large reheating, effect of 1+1?

summary

- Simplified way to compute wall velocities with SM-like friction
- for a moderately strong phase transition, the wall moves subsonic in the dim-6 model

- major uncertainties:

what friction constant to take for very strong phase transitions?

effect of infrared gauge field condensate?

instabilities, turbulence?

aim: use this model to do 3d numerical simulations of the electroweak phase transition (e.g. for GW production)